RECOMMENDATION ITU-R M.1478

PROTECTION CRITERIA FOR COSPAS-SARSAT SEARCH AND RESCUE PROCESSORS IN THE BAND 406-406.1 MHz

(2000)

The ITU Radiocommunication Assembly,

considering

a) that the Cospas-Sarsat global satellite-based search and rescue system operates within an exclusive allocation in the band 406-406.1 MHz;

b) that there is a significant shortfall of spectrum for the non-geostationary (non-GSO) mobile-satellite service (MSS) below 1 GHz, and there is an urgent need to make additional spectrum available on a worldwide basis for such non-GSO MSS systems;

c) that Resolution 219 (WRC-97) invited ITU-R to study the possible use of the band 405-406 MHz by the MSS;

d) that Resolution 219 (WRC-97) resolved to invite ITU-R as a matter of urgency, to study, with the participation of the Inter-Union Commission on Frequency Allocations for Radio Astronomy and Space Science (IUCAF) and other relevant entities, the impact of unwanted emissions on the Cospas-Sarsat system in the band 406-406.1 MHz and the radio astronomy service in the band 406.1-410 MHz and identify appropriate protection measures for these services;

e) the analysis provided in Annex 1 concerning maximum allowable spectral power flux-density (spfd) requirements of the Sarsat Search and Rescue Processor (SARP) against broadband out-of-band emissions and the upper bound on Doppler frequency shift associated with MSS transmissions received by Sarsat;

f) the analysis provided in Annex 2 concerning maximum allowable power flux-density (pfd) requirements for the Sarsat SARP against narrow-band spurious emissions;

g) that Annex 3 provides guidelines for using the protection requirements of the 406-406.1 MHz band for the Sarsat SARP instruments (on-board satellite equipment);

h) that Annex 4 provides protection of the primary safety services allocated in the band 406-406.1 MHz (C-S system) from non-GSO MSS downlink emissions below 406 MHz,

recommends

1 that analysis to determine the effect upon Sarsat SARP instruments by systems using adjacent frequency bands be based upon a maximum acceptable spfd at the Sarsat antenna of $-198.6 \text{ dB}(W/(m^2 \cdot \text{Hz}))$;

2 that analysis to determine the effect upon the Sarsat SARP instruments from narrow-band spurious emissions (e.g. harmonic emissions, parasitic emissions, intermodulation products and frequency conversion products) be based upon a maximum pfd of $-185.8 \text{ dB}(W/m^2)$ at the Sarsat antenna within a resolution bandwidth of 19 Hz;

3 that analysis to determine the effect upon Cospas SARP instruments by systems using adjacent frequency bands be based upon a maximum acceptable spfd at the Cospas antenna of $-200.8 \text{ dB}(W/(m^2 \cdot Hz))$;

4 that analysis to determine the effect upon the Cospas SARP instruments from narrow-band spurious emissions (e.g. harmonic emissions, parasitic emissions, intermodulation products and frequency conversion products) be based upon a maximum pfd of $-176.6 \text{ dB}(W/m^2)$ at the Cospas antenna within a resolution bandwidth of 40 Hz;

5 that analysis to determine the effect upon Cospas-Sarsat non-GSO instruments by proposed MSS systems using the 405-406 MHz frequency band utilize an upper bound Doppler shift of 20 kHz.

ANNEX 1

Protection criteria for Cospas-Sarsat in the band 406-406.1 MHz against out-of-band broadband emissions

1 Introduction

WRC-97 noted the possible use of the band 405-406 MHz as a downlink band by future non-GSO MSS systems and approved Resolution 219 (WRC-97) requesting that analysis be performed to determine the impact this would have on systems using the upper adjacent frequency bands. According to Resolution 219 (WRC-97), any analysis concerning the use of this band by MSS should be limited to systems using narrow-band modulation techniques until further ITU-R studies conclude that other modulation techniques can provide adequate protection to Cospas-Sarsat (406-406.1 MHz) and the radio astronomy services (406.1-410 MHz).

This Annex provides additional information relating to the C-S system and its protection requirements from broadband out-of-band emissions.

2 Background

Other ITU texts provide substantial information concerning the following items:

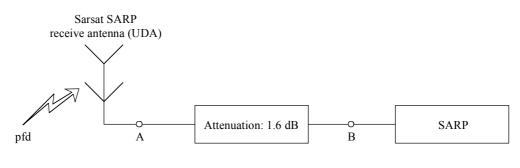
- parameters of several non-GSO MSS networks;
- pfd threshold level of interference;
- search and rescue (SAR) protection using spectral shaping or filtering techniques.

3 spfd threshold level of interference

The addition of broadband noise to the Sarsat SARP will have the effect of increasing the system bit error ratio (BER), and therefore adversely affect its performance. As identified in ITU-R studies the maximum acceptable uplink BER for the Sarsat SARP cannot exceed 5×10^{-5} . Based upon this requirement, this analysis identifies the maximum acceptable pfd associated with broadband noise in the Sarsat SARP uplink channel. The analysis does not address the effect of narrow-band emissions (e.g. spectral lines), which will also adversely affect the SARP's performance nor does it address the protection requirements for all C-S instruments (e.g. Sarsat search and rescue repeater, Cospas SARP).

Figure 1 shows the main hardware elements on board the TIROS satellites (and on the future METOP satellites).

FIGURE 1 On-board hardware equipment



1478-01

The UDA antenna gain pattern specification is expressed according to the nadir angle in Table 1:

TABLE 1

SARP receive antenna (UDA) gain pattern

Nadir satellite angle	62	59	54	47	39	31	22	13	5	0
Gain in RHCP	3.85	3.54	2.62	1.24	-0.17	-1.33	-2.24	-3.08	-3.80	-3.96
Gain in LHCP	-5.69	-6.23	-7.52	-9.39	-11.39	-13.12	-14.52	-15.77	-17.17	-18.00
Axial ratio	6.02	5.85	5.59	5.26	4.90	4.57	4.31	4.11	3.78	3.49

The specified figures in Table 1 are of the 406 MHz-Sarsat receive antenna pattern for the SARP, as they should be for the TIROS and METOP satellites.

The Sarsat typical figures are: noise figure = 2.5 dB (C-S SARP input parameter), nominal background noise temperature = 1 000 K (C-S input parameter), attenuation between the antenna and the SARP receiver = 1.6 dB. Thus, the system noise temperature at the input of the SARP receiver (point B on Fig. 1) equals 1 010 K and therefore, the noise spectral density equals $N_0 = -198.6$ dB(W/Hz).

The worst-case specification states that the SARP is designed to operate correctly when the received signal has a power C = -161 dBW (minimum level of the received signal) at the input of the receiver, which provides an effective $E_b/N_0 = 9.1$ dB in the bit detector of the SARP if we take into account the beacon waveform and the various losses. In this case, the corresponding BER equals 2.6×10^{-5} .

Therefore, in order to achieve a BER of 5×10^{-5} (which is an approximate doubling of the BER) the maximum acceptable degradation is 0.3 dB. At $E_b/N_0 = 8.8$ dB, the BER equals 4.8×10^{-5} .

Hereunder, the additive noise corresponding to the 0.3 dB degradation for the C/N_0 is calculated.

Let I_0 represent the additive noise power density coming from the non-GSO MSS interferers.

The initial N_0 noise becomes $N_0 + I_0$.

The signal-to-noise ratio C/N_0 becomes $C/(N_0 + I_0)$.

The degradation is 0.3 dB = 10 log ((C/N_0)/($C/(N_0 + I_0)$)), thus $I_0/N_0 = -11.5$ dB and $I_0 = -210.1$ dB(W/Hz) which corresponds to a temperature of 70.8 K, and therefore an increase of 7% of the system noise temperature at the input of the SARP receiver.

Therefore, the maximum admissible level of noise density is $I_0 = -210.1 \text{ dB}(\text{W/Hz})$ (calculated for point B in Fig. 1).

As shown in Fig. 1, the noise density, I_0 , takes into account the attenuation and the antenna gain. As the spfd is required, it is necessary to transform this figure in dB(W/(m² · Hz)). The equivalent surface area of an antenna having a gain G is $S = G \frac{\lambda^2}{4\pi}$. Therefore, the corresponding spfd equals -210 + 1.6 (losses) $-10 \log_{10} S = -198.6 \text{ dB}(W/(m^2 \cdot \text{Hz}))$, taking into account the highest satellite nadir angle.

The maximum level of broadband noise interference in the band 406-406.1 MHz shall not exceed $-198.6 \text{ dB}(W/(m^2 \cdot \text{Hz}))$ to protect the Sarsat SARP instrument.

4 Upper bound on Doppler shift

Any proposed protection bandwidth should also account for the Doppler shifts. The value of the maximum Doppler shift must be carefully examined. The worst case occurs when the Sarsat and the non-GSO MSS satellites are located on the same orbit and travel in opposite directions. In this case, the analysis below applies.

The non-GSO MSS signal comes from point A. The Sarsat satellite is represented by point B. The Sarsat satellite is moving at a speed V_B . If the non-GSO MSS satellite is not moving, the received frequency at B is $F_B = F_A \left(1 + \frac{V_B}{C}\right)$ in the worst case. On the other hand, the received frequency at B has the same value if the Sarsat satellite is not moving and if the non-GSO MSS satellite is moving. If the altitude of the satellite equals 850 km, its speed is 7426 km/s.

As the two satellites are moving in opposite directions, the upper bound Doppler shift equals:

$$2F_A(V_B/c) = 20 \text{ kHz}$$

This is a worst-case situation and is not necessarily applicable to all the proposed MSS systems.

5 Conclusions and recommendations

Following the above computations, the conclusions and recommendations regarding the impact of emissions from adjacent frequency bands on the Sarsat SARP are:

- the maximum level of broadband noise interference in the band 406-406.1 MHz shall not exceed $-198.6 \text{ dB}(W/(m^2 \cdot \text{Hz}))$ to protect the Sarsat SARP instrument;
- the upper bound on Doppler shift is 20 kHz;
- it is recommended that further analyses be conducted to determine the impact on C-S from MSS occupying the 405-406 MHz band using a spfd of -198.6 dB(W/(m² · Hz)), an appropriate Doppler shift and accounting for the worst-case scenario associated with the entire MSS constellation as envisaged.

ANNEX 2

Protection criteria for C-S system in the band 406-406.1 MHz against narrow-band spurious emissions

1 Introduction

WRC-97 noted the possible use of the band 405-406 MHz as a downlink band by future non-GSO MSS systems and approved Resolution 219 (WRC-97) requesting analysis be performed to determine the impact this would have on systems using the upper adjacent frequency bands. According to Resolution 219 (WRC-97), any analysis concerning the use of this band by MSS should be limited to systems using narrow-band modulation techniques until further ITU-R studies conclude that other modulation techniques can provide adequate protection to C-S (406-406.1 MHz) and the radio astronomy services (406.1-410 MHz).

This Annex provides information relating to the C-S system and its protection requirements against narrow-band spurious emissions.

2 Background

Annex 1 contains the protection criteria for Sarsat SARP in the band 406-406.1 MHz to be used as a basis for analysis of interference from out-of-band emissions. This Annex provides protection requirements for the Sarsat SARP instrument in respect of interference from narrow-band spurious emissions (harmonic emissions, parasitic emissions, intermodulation products and frequency conversion products).

The terminology used in this Annex is derived from Recommendation ITU-R SM.328 – Spectra and bandwidth of emissions, and from Recommendation ITU-R SM.329 – Spurious emissions.

This Annex addresses protection criteria for only Sarsat SARP instruments and does not necessarily represent the protection criteria for all Cospas-Sarsat instruments.

3 Protection requirement from narrow-band spurious emissions

Figure 1 shows the main Sarsat SARP hardware elements.

To better understand the rationale of this specification, it is necessary to briefly recall the functioning of the SARP instrument.

Sarsat distress beacon transmissions begin with 160 ms of unmodulated carrier to allow a phase-locked loop to lock more easily on the carrier. Figure 2 represents the C-S message format.

FIGURE 2

C-S message format

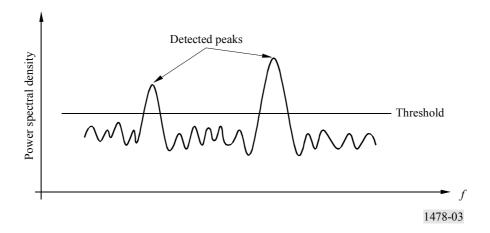
160 ms carrier	Synchronization bits	Distress beacon message content bits
----------------	----------------------	---

1478-02

A spectrum analyser in the SARP instrument continuously monitors the full coverage bandwidth in search of the pure carrier portion of distress beacon transmissions. When the spectrum analyser detects such a line, it considers that it is the beginning of a C-S message. The theory is based on the detection of a pure carrier wave (sine wave) in a white, additive and Gaussian noise environment. The power spectral density of the received signal (pure carrier + noise) is computed using fast Fourier transform techniques, and each signal above the system threshold is processed as if it were a distress beacon (see Fig. 3).

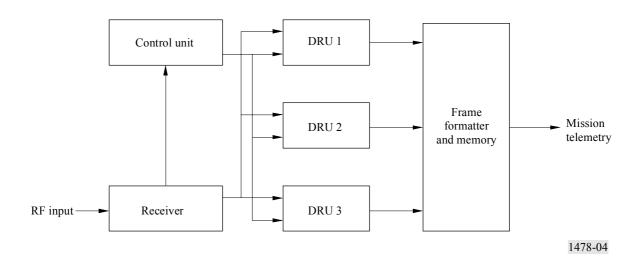
FIGURE 3

Detection of a sine wave in white Gaussian noise



Signals above the threshold level are assigned to an on-board data recovery unit (DRU) for further processing and transmission to the Earth on the mission telemetry channel (see Fig. 4).

FIGURE 4 SARP functional diagram



In order to satisfy SAR performance requirements in respect of low power distress beacons, the Sarsat SARP instrument has been designed to detect and process extremely weak signals. Its performance is such that any signal, C_{min} , which exceeds the local noise density level by 21 dB(Hz) ($C_{min}/N_0 > 21$ dB(Hz)) would be assigned to a DRU for additional processing. Consequently, narrow-band interfering signals meeting this criteria would cause a DRU to be assigned to it. The consequence would be that the performance of the SARP, in terms of capacity (e.g. the number of simultaneous distress messages that are able to be processed), would be seriously degraded.

The Sarsat typical figures are: noise factor = 2.5 dB (C-S SARP typical figure), nominal background noise temperature = 1 000 K (C-S input parameter), attenuation between the antenna and the SARP receiver = 1.6 dB. Thus, the system noise temperature at the input of the SARP receiver (point B on Fig. 1) equals 1 010 K and therefore, the noise spectral density equals $N_0 = -198.6 \text{ dB}(W/\text{Hz})$.

As $C_{min}/N_0 = 21$ dB(Hz), $C_{min} = -177.6$ dBW. Therefore any narrow-band spurious emission greater than -177.6 dBW at the input of the SARP (point B of Fig. 1), would result in a degradation to system capacity.

It is then necessary to compute this maximum admissible level of spectral line at the input of the Sarsat antenna.

The Sarsat SARP receive antenna gain pattern specification is expressed according to the nadir angle in Table 2.

TABLE 2

SARP receive antenna (UDA) gain pattern

Nadir satellite angle	62	59	54	47	39	31	22	13	5	0
Gain in RHCP	3.85	3.54	2.62	1.24	-0.17	-1.33	-2.24	-3.08	-3.80	-3.96
Gain in LHCP	-5.69	-6.23	-7.52	-9.39	-11.39	-13.12	-14.52	-15.77	-17.17	-18.00
Axial ratio	6.02	5.85	5.59	5.26	4.90	4.57	4.31	4.11	3.78	3.49

Therefore, the maximum admissible power at point A of Fig. 1 equals -177.6 + 1.6 (losses) = -176 dBW, taking into account the highest satellite nadir angle. As the pfd is required, it is necessary to transform this figure in dB(W/m²). The equivalent surface area of an antenna having a gain G is $S = G \frac{\lambda^2}{4\pi} = 0.105 \text{ m}^2$ corresponding to the highest satellite nadir angle. Therefore, the corresponding pfd equals $-176 - 10 \log_{10} S = -185.8 \text{ dB}(W/m^2)$.

The required protection level is: no narrow-band spurious emission above $-185.8 \text{ dB}(W/m^2)$ at the input of any Sarsat SARP satellite antenna.

4 Conclusion

Following the above computations, the conclusions and recommendations regarding the impact of spurious narrow-band emissions, on the Sarsat SARP shall not exceed $-185.8 \text{ dB}(W/m^2)$ at the input of any Sarsat SARP antenna.

ANNEX 3

Guidelines for using the protection requirements of the 406-406.1 MHz band (C-S system)

1 Definitions of characteristics of emissions

1.1 Out-of-band emission

Emission on a frequency or frequencies immediately outside the necessary bandwidth which results from the modulation process, but excluding spurious emissions.

1.2 Spurious emission

Emission on a frequency, or frequencies, which are outside the necessary bandwidth and the level of which may be reduced without affecting the corresponding transmission of information. Spurious emissions include harmonic emissions, parasitic emissions, intermodulation products and frequency conversion products, but exclude out-of-band emissions.

1.3 Unwanted emissions

Consist of spurious emissions and out-of-band emissions.

1.4 Necessary bandwidth

For a given class of emission, the width of the frequency band which is just sufficient to ensure the transmission of information at the rate and with the quality required under specified conditions.

2 Procedure for computing the level of unwanted emissions relating to the Sarsat SARP

2.1 Out-of-band emissions

The unit of an unwanted out-of-band emission is $dB(W/(m^2 \cdot Hz))$ (spfd). The overall spfd is in fact an aggregate spfd which is defined as the summation of all spfd coming from the whole potential sources of unwanted out-of-band emissions.

2.2 Spurious emissions

The C-S receiver processors are designed to detect discrete spectral components (unmodulated beacon carrier). The protection requirement is expressed in terms of pfd, and the unit of an spurious emission is $dB(W/m^2)$ (pfd).

The resolution bandwidth of the receiver of the Sarsat SARP instrument is 19 Hz. It means that the minimum frequency spacing between spectral lines (also called resolving power of the spectrum analysis) that the Sarsat SARP receiver may detect, equals 19 Hz. Consequently, a spurious emission level should be computed within a reference bandwidth of 19 Hz.

2.3 Unwanted emission limits

Figure 5 recaps the values of unwanted emission limits.

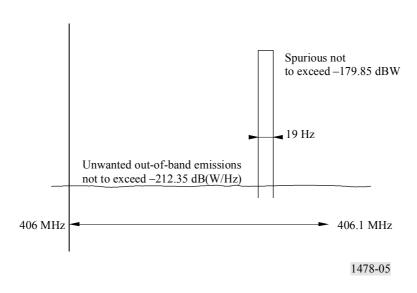


FIGURE 5 Levels of unwanted emissions

The unwanted out-of-band emissions must not exceed $-198.6 \text{ dB}(W/(m^2 \cdot \text{Hz}))$ at the input of any Sarsat SARP antenna: this figure can also be transformed into a spectral power density (dB(W/Hz)):

-210.1 dB(W/Hz) (at the input of the SARP instrument) + 1.6 (losses) -3.85 (antenna gain) = -212.35 dB(W/Hz) at the input of any Sarsat SARP antenna.

The narrow-band spurious emission must not exceed $-185.8 \text{ dB}(\text{W/m}^2)$ at the input of any Sarsat satellite antenna: this figure can also be transformed into a power level (dBW):

-177.6 dBW (at the input of the SARP instrument) + 1.6 (losses) -3.85 (antenna gain) = -179.85 dBW at the input of any Sarsat SARP antenna.

All the above values are valid at the input of any Sarsat SARP antenna.

ANNEX 4

Protection of the primary safety services allocated in the band 406-406.1 MHz (C-S system) from non-GSO MSS downlink emissions below 406 MHz

1 Introduction

WRC-97 noted the possible use of the band 405-406 MHz as a downlink band by future non-GSO MSS systems. Following Resolution 219 (WRC-97), the use should be limited to systems using narrow-band modulation techniques until further ITU-R studies conclude that other modulation techniques can protect C-S (406-406.1 MHz) and the radio astronomy services (406.1-410 MHz).

The international C-S system has been operational since 1985 and has already contributed to saving thousands of lives worldwide. It is, therefore, essential to protect its proper operation.

2 Background

Other ITU texts provide substantial information regarding the following items:

- parameters of several non-GSO MSS networks;
- pfd threshold level of interference; and
- SAR protection using spectral shaping or filtering techniques.

Annex 1 provides protection criteria for Sarsat SARP in the band 406-406.1 MHz. This Annex provides additional information relating to Cospas SARP receiver protection requirements.

3 Protection criteria for Cospas SARP in the band 406-406.1 MHz against out-of-band broadband emissions

The performance of the SARP is expressed by its BER and is directly related to the signal-to-noise density ratio, C/N_0 (dB(Hz)).

The Cospas receiver noise temperature is 300 K. The nominal background noise temperature varies from 300 K in Antarctica to 1000 K in highly populated areas. As a distress beacon signal can be originated from any point on the globe and its decoding should not be impaired by interference from MSS satellite, the most interference sensitive case occurs at the lowest background noise temperature i.e. 300 K.

The attenuation between the antenna and the Cospas SARP receiver is 1.6 dB. The cable noise temperature is 300 K. Thus, the system noise temperature at the input of the SARP receiver equals 600 K and, therefore, the noise spectral density, N_0 is -200.82 dB(W/Hz).

The specified Cospas SARP BER is less than 1×10^{-5} . In accordance with a theoretical BPSK modulation performance, this corresponds to $E_b/N_0 = 9.6$ dB in the bit detector of the SARP.

As identified in ITU-R studies the maximum acceptable BER of the Cospas-Sarsat uplink should not exceed 5×10^{-5} which is achieved when $E_b/N_0 = 9.1$ dB.

Therefore, the maximum acceptable degradation to the E_b/N_0 is 0.5 dB, which corresponds to increase of noise density $N_0 = 0.5$ dB.

If *I* is the additive noise power density contribution from non-GSO MSS interferers, then the value of noise density becomes $N_0 + I$, and the bit energy to noise density ratio becomes $E_b/(N_0 + I)$.

The acceptable degradation is 0.5 dB = 10 log $((E_b/N_0)/(E_b/(N_0 + I)))$. Thus $I/N_0 = -9.14$ dB and I = -209.95 dB(W/Hz) which corresponds to a noise temperature increase of 73.21 K (an increase of 12% of the system noise temperature at the input of the Cospas SARP receiver).

The maximum admissible level of noise density, I, is -209.95 dB(W/Hz) at the input satellite receiver.

The noise density, *I*, takes into account the attenuation and the antenna gain. To express the figure in spfd, it is necessary to transform this value to dB(W/(m² · Hz)). The equivalent surface area of an antenna having a gain *G* is $S = G \frac{\lambda^2}{4\pi}$. As the maximum gain of the spiral cone antenna installed on the Cospas spacecraft is 6 dB, then S = 0.174 m². Therefore, the corresponding spfd equals -209.95 + 1.6 (losses) $- 10 \log_{10} S = -200.8$ dB(W/(m² · Hz)).

The maximum level of broadband noise interference in the band 406-406.1 MHz shall not exceed $-200.8 \text{ dB}(W/(m^2 \cdot \text{Hz}))$ in order to provide adequate protection for the Cospas SARP instrument.

4 Protection criteria for Cospas SARP in the band 406-406.1 MHz against narrow-band spurious emissions

The Cospas SARP continuously scans the 406-406.1 MHz band (one scan duration is 60 ms) in search of beacon signals which exceed the SARP detection threshold. After a signal is detected the SARP DRU will be locked for at least 520 ms on the signal frequency. A single interfering signal from an MSS satellite in the 406-406.1 MHz band could capture both on board DRUs making a Cospas satellite unavailable for processing real beacon distress signals.

In order to satisfy SAR performance requirements in respect of low power distress beacons, the Cospas SARP instrument has been designed to detect and process extremely weak signals. Its performance is such that any emission (C_{min}) at a power level which exceeds the power noise density level by 21 dB ($C_{min}/N_0 > 21$ dB(Hz)) would be assigned to a DRU for additional processing. Consequently, any interfering signals in the processed bandwidth 406-406.1 MHz meeting this criteria would cause a DRU to be assigned to it. Therefore, a beacon message would be lost even if the narrow-band interfering signal did not affect the beacon signal directly.

The noise spectral density, N_0 , was calculated to be -200.82 dB(W/Hz).

As $C_{min}/N_0 = 21$ dB(Hz), $C_{min} = -179.82$ dBW. Therefore, any narrow-band spurious emission greater than -179.82 dBW at the input of the SARP would degrade the Cospas SARP beacon processing capacity.

The Cospas SARP receive antenna maximum gain is 6 dB. Therefore, the maximum admissible power at the antenna input is equal to -179.82 + 1.6 (losses) = -178.22 dBW. As the pfd is required, it is necessary to transform this figure to dB(W/m²). The equivalent surface area of an antenna having the gain G = 6 dB is $S = G \frac{\lambda^2}{4\pi} = 0.174$ m². Therefore, the corresponding flux-density equals $-178.2 - 10 \log_{10} S = -176.6$ dB(W/m²).

The maximum level of narrow-band spurious emissions should not exceed $-176.6 \text{ dB}(W/m^2)$ at the input of Cospas SARP satellite antenna in order to provide adequate protection for the Cospas SARP instrument.

The C-S receiver processors are designed to detect discrete spectral components (unmodulated beacon carrier). Therefore, the protection requirement is expressed in terms of power spectrum rather then power spectrum density, and the unit of spurious emission is $dB(W/m^2)$ (pfd). The resolution bandwidth of the receiver of the Cospas SARP instrument is 40 Hz. Therefore, the minimum frequency spacing between spectral lines (also called resolving power of the spectrum analysis) that the Cospas SARP receiver can detect is 40 Hz. Consequently, a spurious emission level should be computed within a reference bandwidth of 40 Hz.

5 Desensitization and front-end burn out of the Cospas receiver

The Cospas satellites use high sensitivity receivers with wideband low noise amplifiers (LNAs) and could experience overload or, in the worst case, front-end burn out from non-GSO MSS emissions.

Blocking or desensitization are generally observed when the lack of RF filtering cause the saturation of the front stage LNA and thus a decrease in its gain and in the receiver sensitivity. Other effects such as the mixing of the interfering signal with the phase noise of the LNA can combine with the above. As a general observation, the experience shows that when a narrow-band transmitter operates close to a wideband receiver, the blocking or desensitization is often the dominant cause of interference problems.

Cospas 406 MHz receiver LNAs can withstand signals of -60 dBW (front-end burn out) and -100 dBW (desensitization). As the non-GSO MSS satellite power can be as high as 24 dBW it is clear that burn-out of Cospas satellite receivers could occur if not prevented by frequency separation.

The power level of non-GSO MSS signals at Cospas receiver input depends on many factors, the most important of which is the distance between the non-GSO MSS and Cospas satellites, which theoretically can be as small as zero for some of the anticipated non-GSO MSS satellite constellations. In view of the catastrophic consequences of the front-end burn out of Cospas receivers, the non-GSO MSS carrier frequencies should be selected to ensure that none of their declared bandwidth extends above 406 MHz. Table 3 presents the key factors for calculating the minimum guardband to protect Cospas receivers against burn out from non-GSO MSS satellite emissions.

TABLE 3

Doppler shift defined	20 kHz 2 × 406 000 × (7.4/300 000)
Stability of Tx of space station for this band	$\begin{array}{c} 8.12 \text{ kHz} \\ 20 \times 10^{-6} \times 406000 \end{array}$
Stability of SARP receiver per year (in assumption of lifetime of satellite 12 years (Sarsat-3 launched in 1986)	2.45 kHz
Non-GSO MSS declared bandwidth/2	1.2-500 kHz (2.4-1 000)/2
Guardband	31.8-530.6 kHz

The results presented in Table 3 indicate that, as a minimum, a 32 kHz guardband (405.968-406 MHz) would be required if the 405-406 MHz frequency band was allocated for use by non-GSO MSS.